

# Validity of the ICFT $R$ -matrix method: Be-like $\text{Al}^{9+}$ a case of study

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- 1 Motivation
- 2 Atomic structure
- 3 Effective collision strengths
- 4 Conclusions

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K. M. Aggarwal, F. P. Keenan 2015, *Mon. Not. R. Astr. Soc.*, **447**, 3849.
- Reply: L. Fernández-Mencherero, G. Del Zanna and N. R. Badnell 2015, *Mon. Not. R. Astr. Soc.*, **450**, 4174.

# Motivation

- 238-level ICFT of MZB-2014 gives in general larger results for the effective collision strengths  $\Upsilon$  than 98-level DARC of AK-2014.
- Electron exchange included by AK-2014 for more angular momenta than MZB-2015:  $2J = 81$  vs  $2J = 23$ .
- Energy step length of AK-2014 slight finer than MZB-2014, depends on ion.
- AK-2014 goes further in scattering energy than MZB-2014: 12 times versus 3 times the ionization potential plus interpolation up to  $E = \infty$  in Burgess–Tully domain.
- DARC vs ICFT.



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- DARC vs ICFT.
- MZB-2014 has a much larger configuration interaction and close coupling expansion than AK-2014: 238 levels vs 98 levels.

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# Atomic energies

**238-level CI structure<sup>1</sup>** configurations  $1s^2 \{2s^2, 2s2p, 2p^2\}, 1s^2 2\{s, p\} nl$ ,  
with  $n = 3 - 7$  and  $l = s, p, d, f, g$  for  $n = 3 - 5$  and  
 $l = s, p, d$  for  $n = 6 - 7$ .  
Total 130 terms, 238 levels.

**98-level CI structure<sup>2</sup>** configurations  $1s^2 \{2s^2, 2s2p, 2p^2\}, 1s^2 2\{s, p\} nl$ ,  
with  $n = 3 - 4$  and  $l = s, p, d, f$ .  
Total 54 terms, 98 levels.

<sup>1</sup> L. Fernández-Menchero, G. Del Zanna and N. R. Badnell 2014, *Astron. Astrophys.*, **566**, A104.

<sup>2</sup> Present work (for test comparison).

# Atomic energies

Table: Al<sup>9+</sup> target levels. All energies are in cm<sup>-1</sup>.

<i>i</i>	Conf. Level	$E_{\text{NIST}}$	$E_{98}$ ( % )	$E_{238}$ ( % )
1	2s <sup>2</sup> <sup>1</sup> S <sub>0</sub>	0.	0. ( 0 )	0. ( 0 )
2	2s 2p <sup>3</sup> P <sub>0</sub> <sup>o</sup>	155148.	155722. (0.4)	155539. (0.3)
3	2s 2p <sup>3</sup> P <sub>1</sub> <sup>o</sup>	156798.	157487. (0.4)	157404. (0.4)
4	2s 2p <sup>3</sup> P <sub>2</sub> <sup>o</sup>	160429.	161146. (0.4)	161278. (0.5)
5	2s 2p <sup>1</sup> P <sub>1</sub> <sup>o</sup>	300490.	309273. (2.9)	307209. (2.2)
6	2p <sup>2</sup> <sup>3</sup> P <sub>0</sub>	404574.	408026. (0.9)	407826. (0.8)
7	2p <sup>2</sup> <sup>3</sup> P <sub>1</sub>	406517.	409969. (0.8)	409888. (0.8)
8	2p <sup>2</sup> <sup>3</sup> P <sub>2</sub>	409690.	413420. (0.9)	413526. (0.9)
9	2p <sup>2</sup> <sup>1</sup> D <sub>2</sub>	449732.	458157. (1.9)	457831. (1.8)
10	2p <sup>2</sup> <sup>1</sup> S <sub>0</sub>	553783.	567794. (2.5)	567267. (2.4)

# Atomic energies

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<i>i</i>	Conf. Level	$E_{\text{NIST}}$	$E_{98}$ (%)	$E_{238}$ (%)
88	2p 4f <sup>3</sup> P <sub>3</sub>	—	2712961. (—)	2706897. (—)
89	2p 4f <sup>3</sup> P <sub>4</sub>	—	2713531. (—)	2707350. (—)
90	2p 4f <sup>3</sup> P <sub>5</sub>	—	2715225. (—)	2708632. (—)
91	2p 4f <sup>3</sup> D <sub>3</sub>	—	2716473. (—)	2710743. (—)
92	2p 4f <sup>1</sup> D <sub>4</sub>	—	2717105. (—)	2710218. (—)
93	2p 4f <sup>3</sup> D <sub>2</sub>	—	2717141. (—)	2711361. (—)
94	2p 4p <sup>1</sup> S <sub>0</sub>	—	2717922. (—)	2706869. (—)
95	2p 4f <sup>3</sup> D <sub>1</sub>	—	2718397. (—)	2712736. (—)
96	2p 4f <sup>1</sup> D <sub>2</sub>	—	2719895. (—)	2713991. (—)
97	2p 4d <sup>1</sup> F <sub>3</sub> <sup>o</sup>	—	2722212. (—)	2715080. (—)
98	2p 4d <sup>1</sup> P <sub>1</sub> <sup>o</sup>	—	2723448. (—)	2716711. (—)

# Infinite energy point

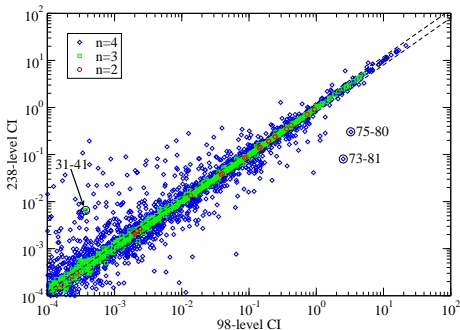


Figure: Transition infinite energy points  $y_\infty = \Omega_\infty^{\text{PWB}}$  or  $\frac{4}{3}S$  for two different atomic structures: 98-level CI versus 238-level CI.

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## 1 Motivation

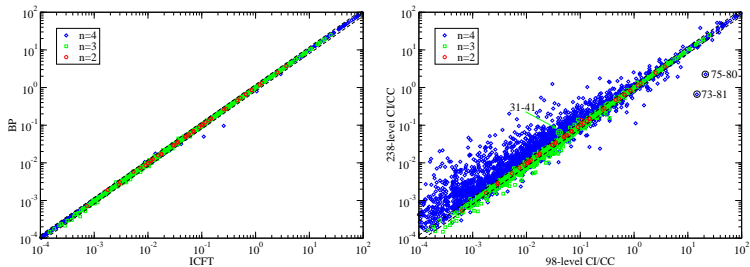
## 2 Atomic structure

## 3 Effective collision strengths

- Peak abundance temperature
- Low temperature
- High temperature

## 4 Conclusions

# Result comparison



**Figure:** Effective collision strengths for electron-impact excitation of the ion  $\text{Al}^{9+}$  at  $T = 10^6$  K. Left: Breit–Pauli versus ICFT, 98 level CI/CC; right: 98-level CI/CC versus 238-level CI/CC, ICFT.

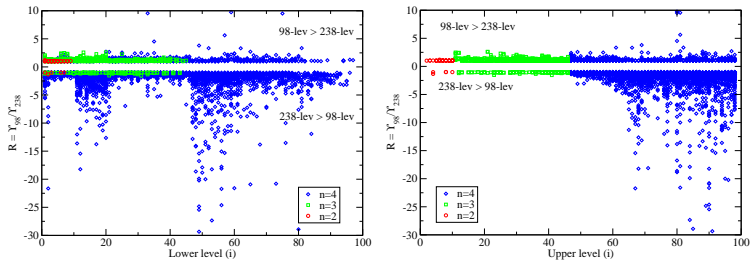


# Result comparison

**Table:** Number of transitions in figures which differ by more than a certain relative error  $\delta = |\Upsilon_A - \Upsilon_B|/\Upsilon_B$  (or  $\Upsilon \rightarrow y_\infty$ ), as a percentage.

Rel. error (%)	$y_\infty$ 98 vs 238 CI	98 CI/CC $\Upsilon$ BP vs ICFT	ICFT $\Upsilon$ 98 vs 238 CI/CC
1	3778	1336	4579
2	3600	803	4400
3	3416	500	4243
4	3266	350	4077
5	3141	260	3928
6	3022	206	3798
7	2914	158	3676
8	2804	127	3569
9	2722	106	3460
10	2644	82	3357
20	2068	22	2582
30	1643	9	2090
40	1356	4	1725
50	1163	2	1449
75	846	2	1113
100	707	2	901
150	538	1	647
200	443	0	505
300	336	0	320
1000	187	0	88
Total	4035	4753	4753

# Result comparison

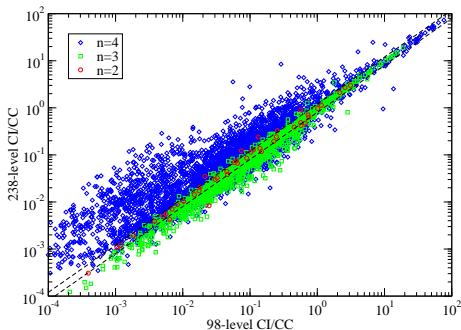


**Figure:** The ratio of effective collision strengths from 98- vs 238-level CI/CC ICFT  $R$ -matrix calculations vs left: lower level and right: upper level index for all inelastic transitions amongst the 98 lowest common levels of  $\text{Al}^{19+}$  at  $T = 10^6 \text{ K}$ . Positive values indicate  $\Upsilon_{98} > \Upsilon_{238}$  and negative values  $\Upsilon_{238} > \Upsilon_{98}$ .

# Result comparison

- Comparing different relativistic formalism the differences are minimal. Only a few spin-change transitions present differences, and they are much smaller compared with CI/CC differences.
- Comparing different atomic structures, differences in the results  $\Upsilon$  are same order as differences in the atomic structure  $\Omega_{\infty}^{\text{PWB}}, \frac{4}{3}S$ .
- $\Upsilon$  for 238-level CI/CC expansion are in general larger than the ones for 98-level CI/CC expansion, while  $\Omega_{\infty}^{\text{PWB}}, \frac{4}{3}S$  was symmetrical.
- Differences are considerable larger for transitions to the  $n = 4$  shell than for the  $n = 2, 3$  ones.
- 98-level CC expansion is not converged for the  $n = 4$  atomic shell.
- 238-level CI expansion gives a much more accurate atomic structure than 98-level.

# Result comparison



**Figure:** Effective collision strengths for electron-impact excitation of the ion  $\text{Al}^{19+}$  at  $T = 2 \times 10^4 \text{ K}$ . 98-level CI/CC versus 238-level CI/CC, ICFT.

# Result comparison

- Differences between atomic structures much larger than at the peak abundance temperature.
- Position or resolution of resonances.
- Both same step lengths and checked convergence. Resolution of the resonances discarded.
- Differences in position of the resonances due to atomic structure.
- 238-level CI expansion much more reliable than 98-level one, mainly for levels in the  $n = 4$  atomic shell.

# Result comparison

**Table:** Comparison of effective collision strengths,  $\Upsilon$ , at  $T = 2 \times 10^7$  K and the reduced quantity  $y_\infty$  at infinite temperature.  $A(B)$  denotes  $A \times 10^B$ .

$i-j$	DARC	ICFT			
	98-level CI/CC	98-level CI/CC		238-level CI/CC	
$T$ (K)	$2 \times 10^7$	$2 \times 10^7$	$\infty$	$2 \times 10^7$	$\infty$
1-52	3.01 (-2)	3.20 (-2)	2.40 (-2)	3.22 (-2)	2.64 (-2)
1-64	1.70 (-4)	1.96 (-4)	1.01 (-4)	9.40 (-5)	1.63 (-5)
1-70	5.29 (-6)	4.23 (-6)	2.68 (-6)	2.18 (-5)	2.18 (-5)
1-80	8.45 (-6)	7.05 (-6)	—	1.09 (-5)	—

DARC: K. M. Aggarwal and F. P. Keenan 2014, *Mon. Not. R. Astr. Soc.*, **438**, 1223

98-level ICFT: present work

238-level ICFT: L. Fernández-Menchero, G. Del Zanna, N. R. Badnell 2014, *Astron. Astrophys.*, **566**, A104.

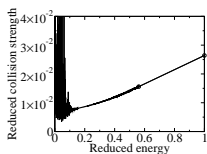
# Differences at high temperature

MZB-2014 calculated  $\Omega$  up to **three** times the ionization potential and interpolated in Burgess–Tully domain.

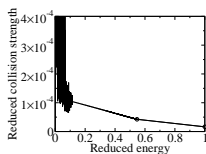
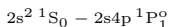
A. Burgess and J. A. Tully 1992. *Astron. Astrophys.*, **254**, 436–453.

AK-2014 calculated  $\Omega$  up to **twelve** times the ionization potential and did not perform any interpolation up to infinity.

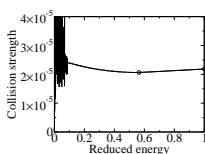
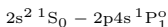
## Burgess–Tully diagram



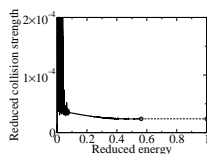
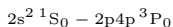
1 – 52



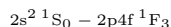
1 – 64



1 – 70



1 – 80



**Figure:** Reduced collision strengths versus energy from 238-level CI/CC ICFT R-matrix calculations for selected transitions of  $\text{Al}^{9+}$ .



# Result comparison

- Calculated  $\Omega$  reached asymptotic form. Interpolation in Burgess–Tully domain properly done.
- Differences in the infinite energy point lead to differences in the  $\Upsilon$  for high temperature.
- Infinite energy points more reliable for 238-level CI expansion than for 98-level CI expansion.

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# Conclusions

- Three main error sources: atomic structure, close coupling expansion and inclusion of relativistic effects.
- Different atomic structures / CI expansions lead to large differences, mainly at high temperature / energy.
- A larger CC expansion improves strongly the results, mainly for the higher excited terms / levels.
- Method to consider the relativistic effects (ICFT, BP, DARCF) does not change meaningfully the results.
- No calculation can guarantee the quality of the results for the last terms / levels included in the close-coupling expansion.
- Nothing special about Be-like  $Al^{9+}$ . Same conclusions can be applied to other ions and sequences.

For example:  $Fe^{13+}$ : G. Del Zanna, N. R. Badnell, L.

Fernández-Mencheró, G. Y. Liang, H. E. Mason and P. J. Storey

2015. *Mon. Not. R. Astron. Soc. In press.*

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## Validity of the ICFT $R$ -matrix method: Be-like Al<sup>9+</sup> a case study

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# Thank you for your attention